

Santa Fe Workshop 2012

July, 2012

**125 GEV HIGGS BOSON, ENHANCED DI-
PHOTON RATE, AND
GAUGED $U(1)_{PQ}$ -EXTENDED MSSM**



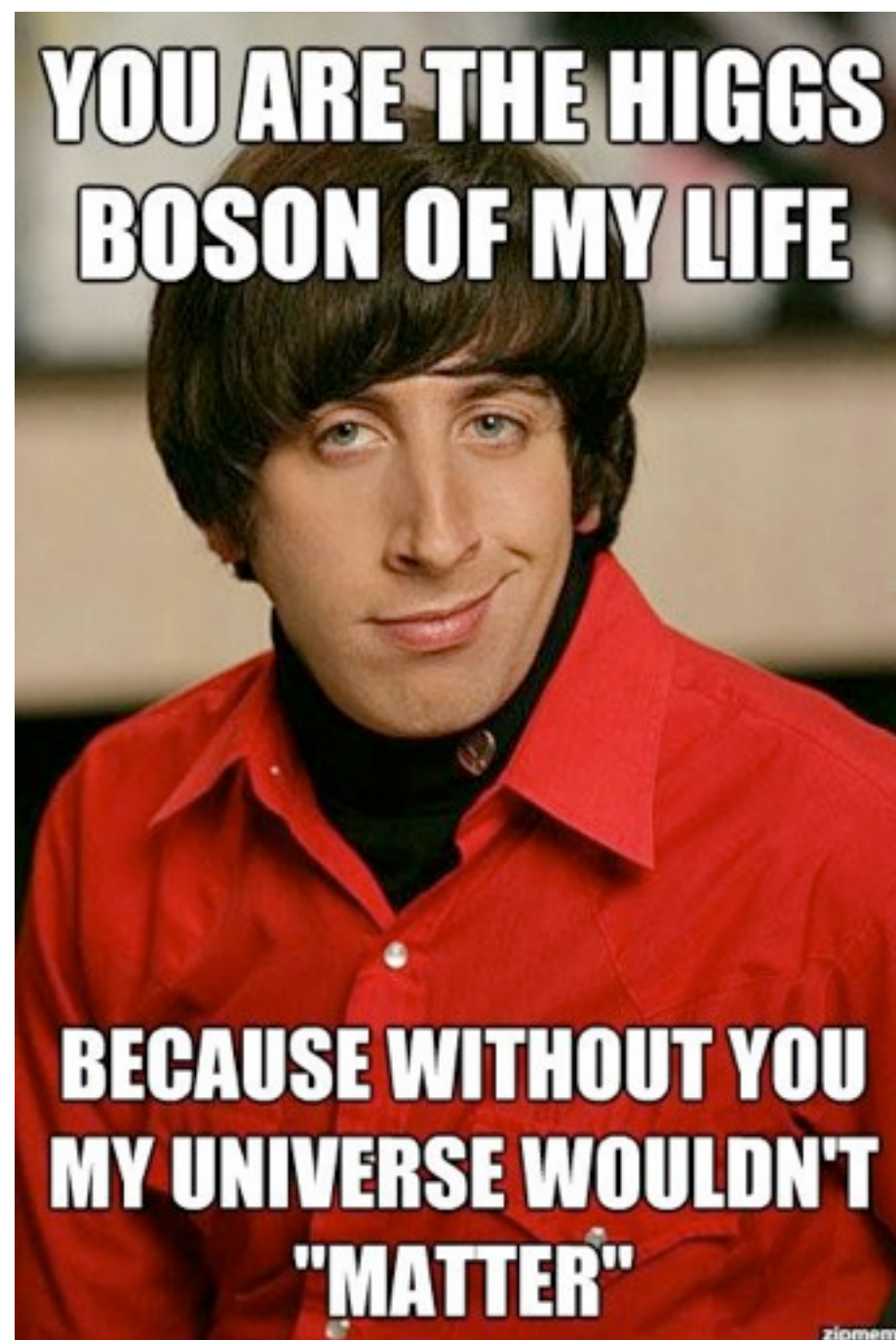
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**Based on Arxiv:1207.2473., in collaboration
with H.P. An and L.T. Wang**



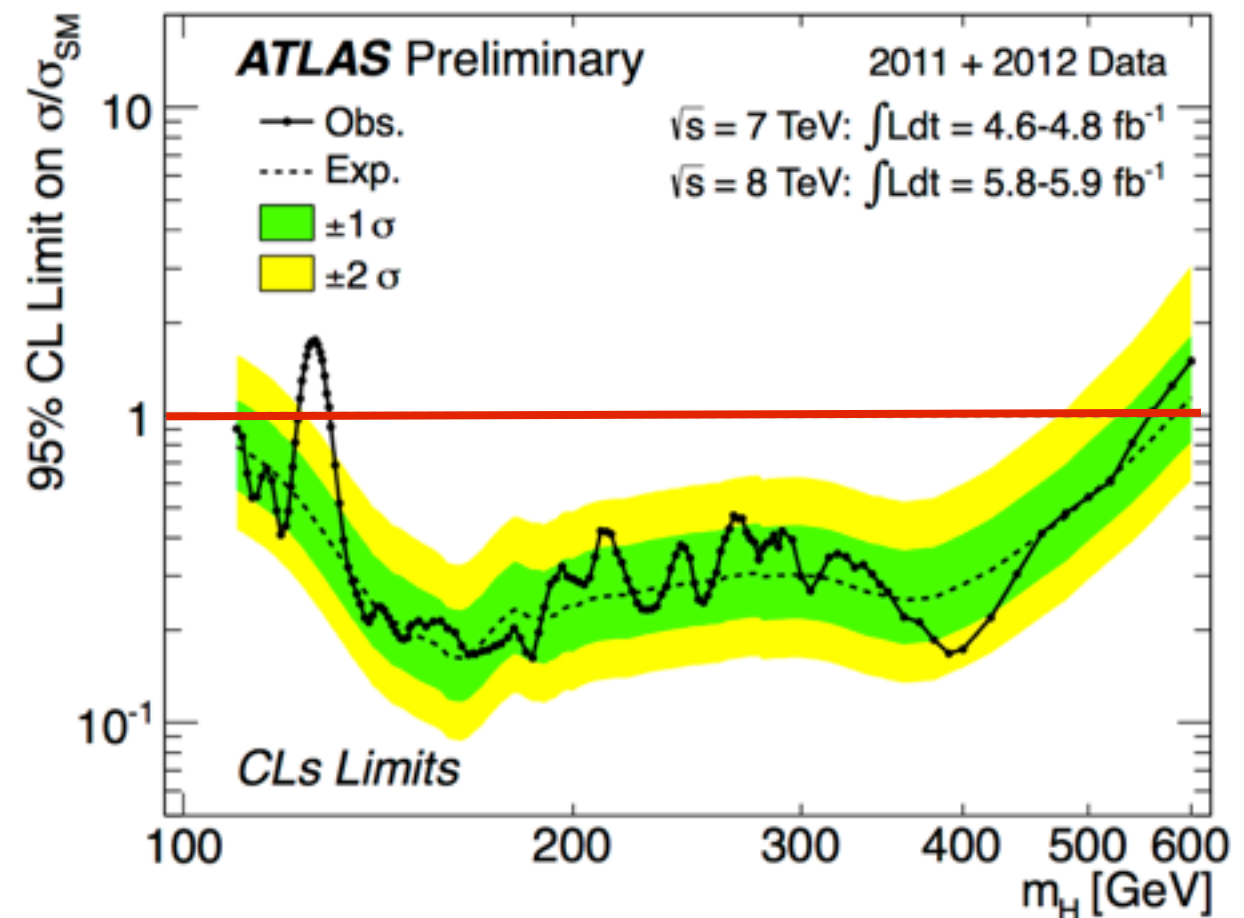
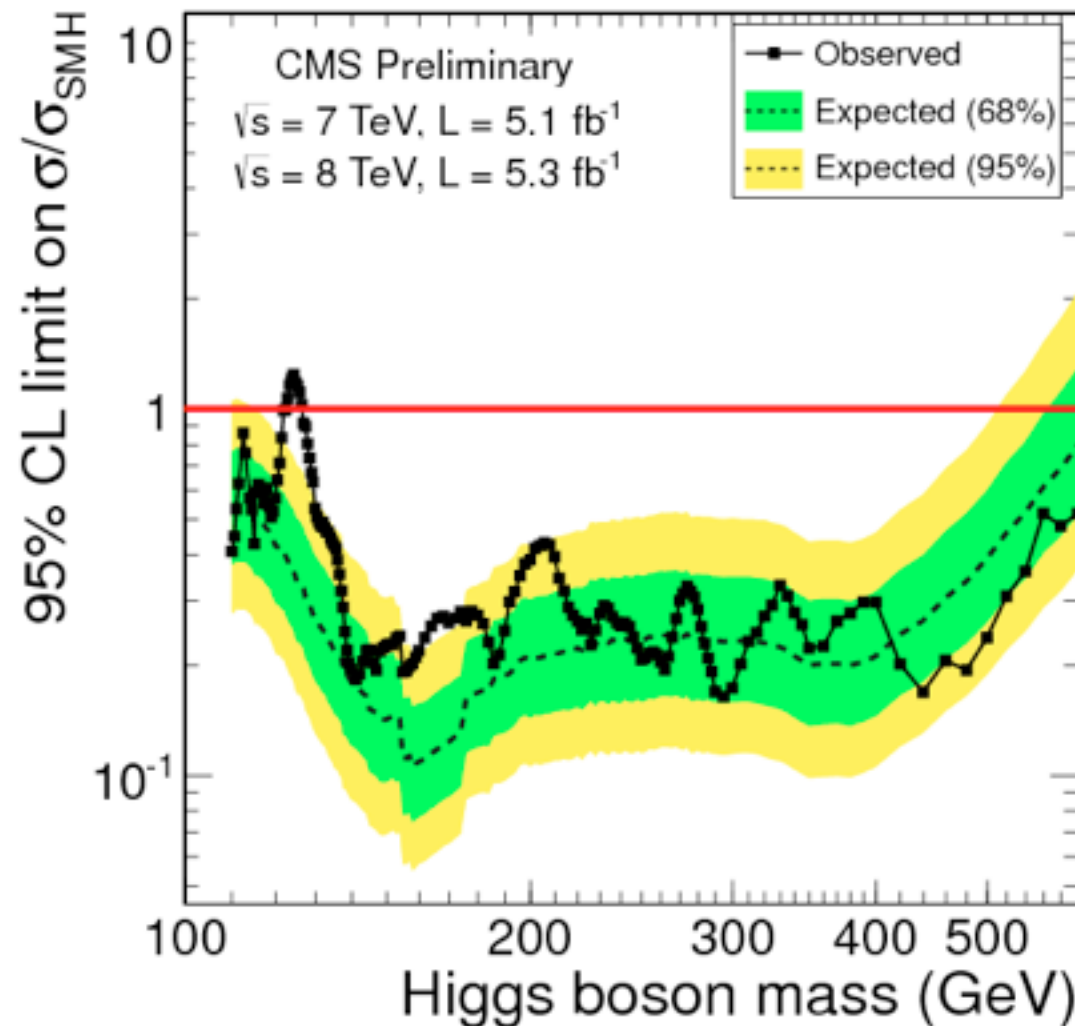
Why Higgs Boson ?





Progress at the LHC

[Atlas+CMS, 2012]

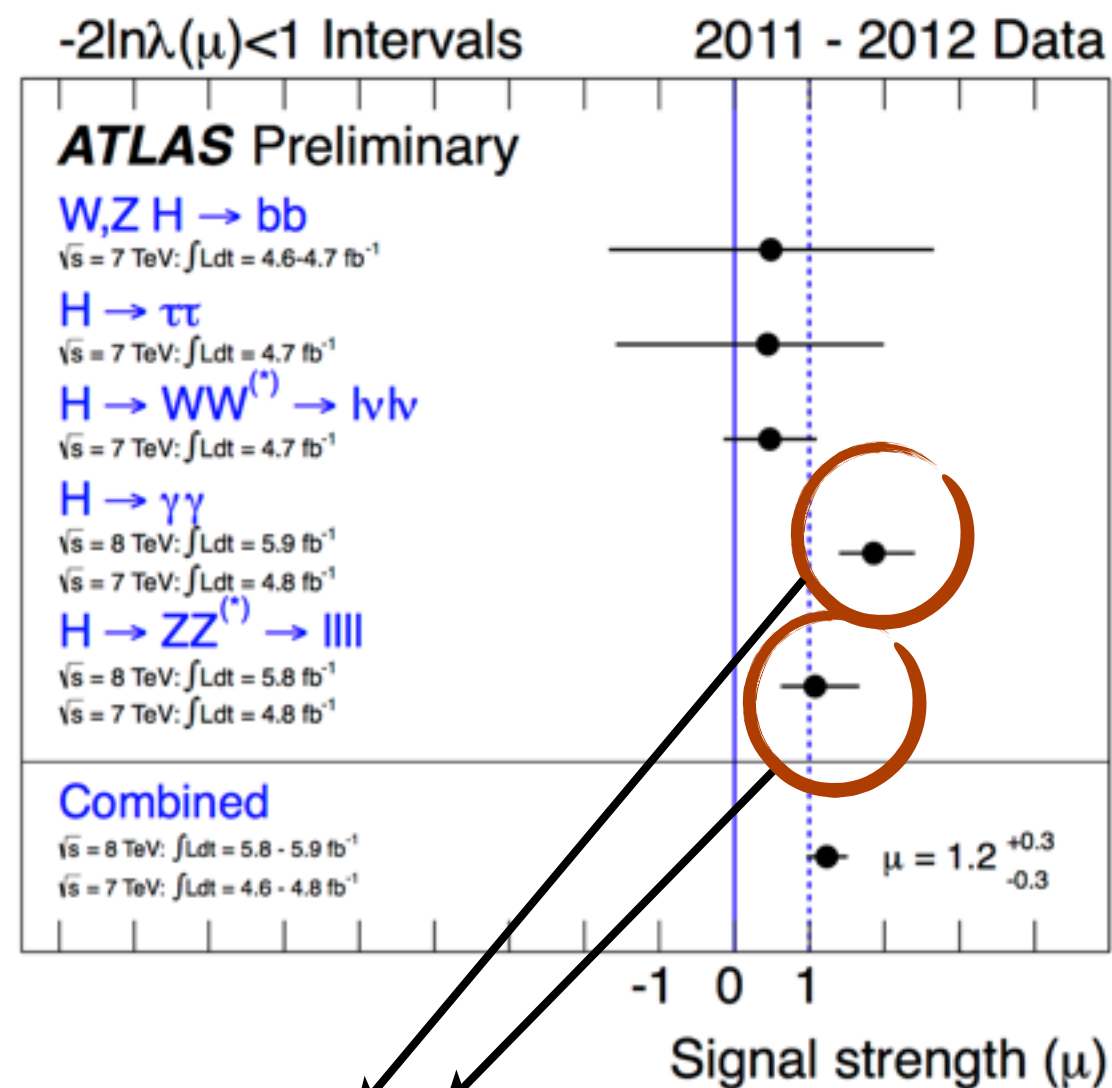
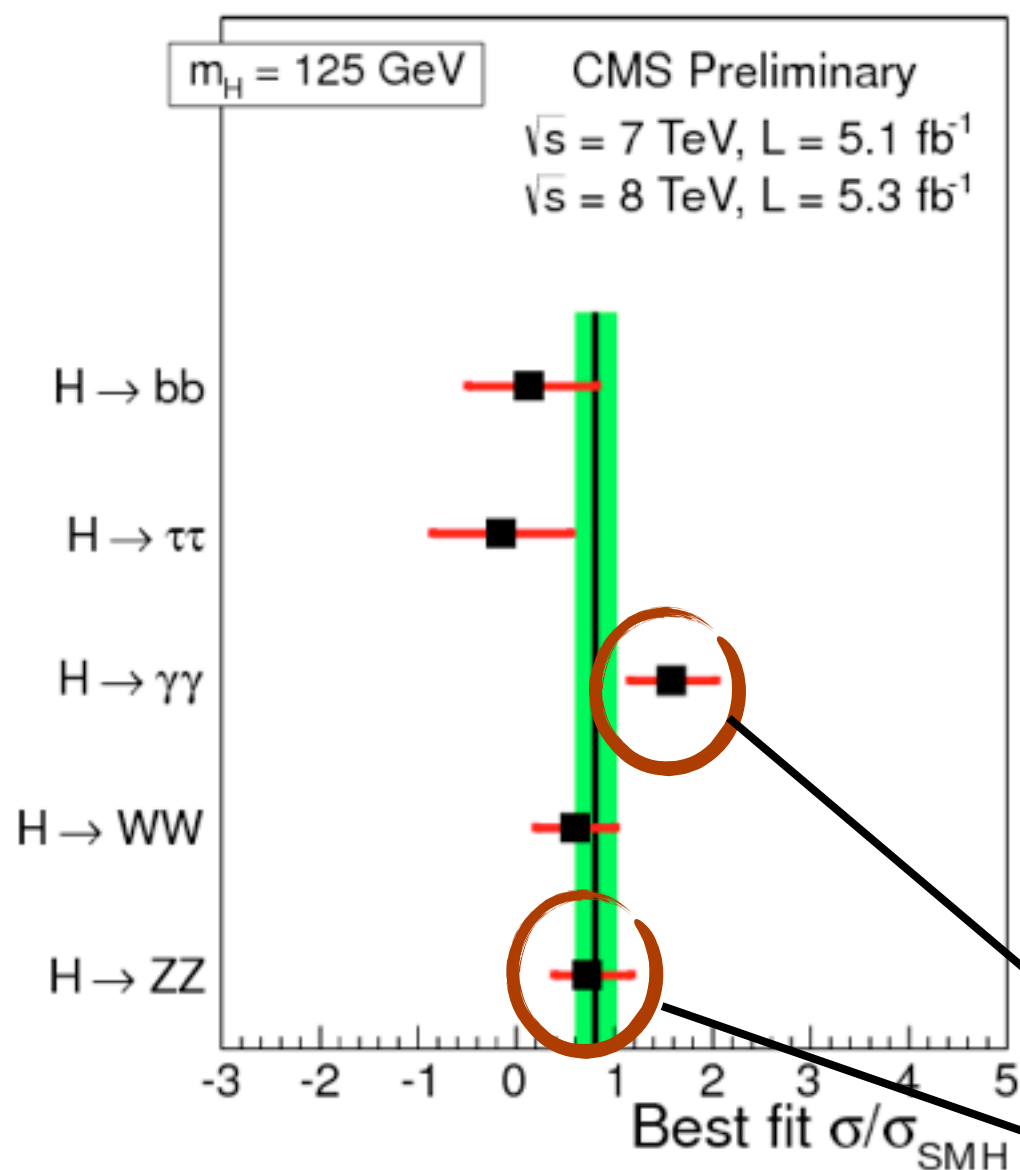


- ☒ CMS: 110-123 ... 127-600 GeV, excluded at 95% C.L.
- ☒ ATLAS: 110-123 ... 130-558 GeV, excluded at 95% C.L.
- ☒ Event excesses are observed where $\sim 125-126 \text{ GeV}$, mainly contributed by di-photon and ZZ searches
- ☒ Under background-only hypothesis, \Rightarrow discovery of a new boson



Can The LHC Results Be Fit Well In Supersymmetry ?

[Atlas + CMS, 2012]



Key quantities for fit: (1) invariant mass $m_h \sim 125 \text{ GeV}$;
(2) signal rates of di-photon and ZZ

A little over interpreting ! But, it is fun to see what it might mean if this is true.



Suppressed $b\bar{b}$ Decay Width?

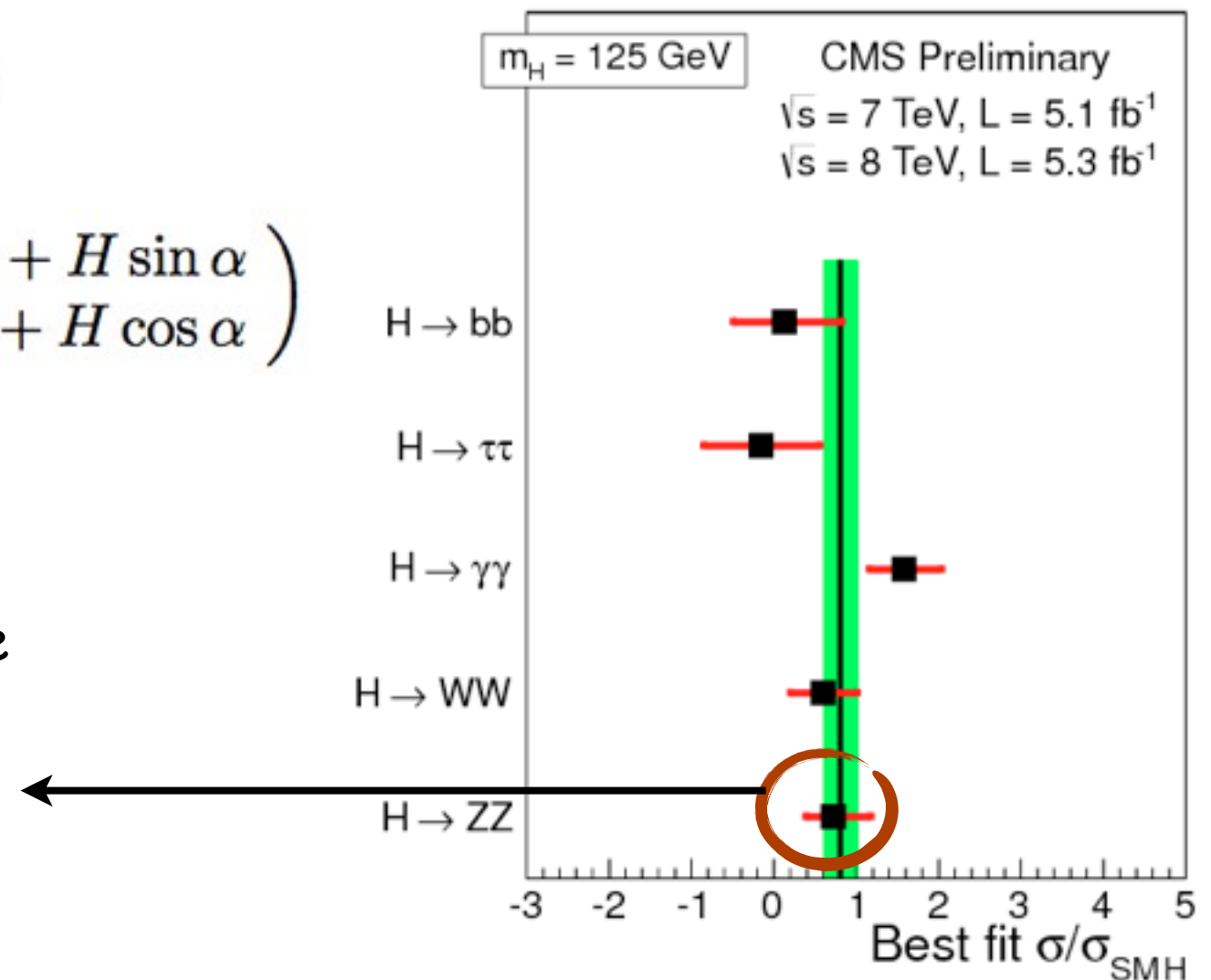
$$\sigma_{\gamma\gamma} = \sigma_{h_{\text{SM}}} \times \frac{\Gamma_{\gamma\gamma}}{\Gamma_{b\bar{b}} + \sum_{\text{SM}}^{i \neq b\bar{b}} \Gamma_i}$$

- Obviously, the di-photon signal rate can be enhanced. (See Spencer's talk)
- In SUSY, this can be achieved by suppressing

$$\frac{y_{hb\bar{b}}}{y_{hb\bar{b}}^{\text{SM}}} = -\frac{\sin \alpha}{\cos \beta}$$

$$\begin{pmatrix} \text{Re}(H_u) \\ \text{Re}(H_d) \end{pmatrix} = \frac{1}{\sqrt{2}} \begin{pmatrix} v_u + h \cos \alpha + H \sin \alpha \\ v_d - h \sin \alpha + H \cos \alpha \end{pmatrix}$$

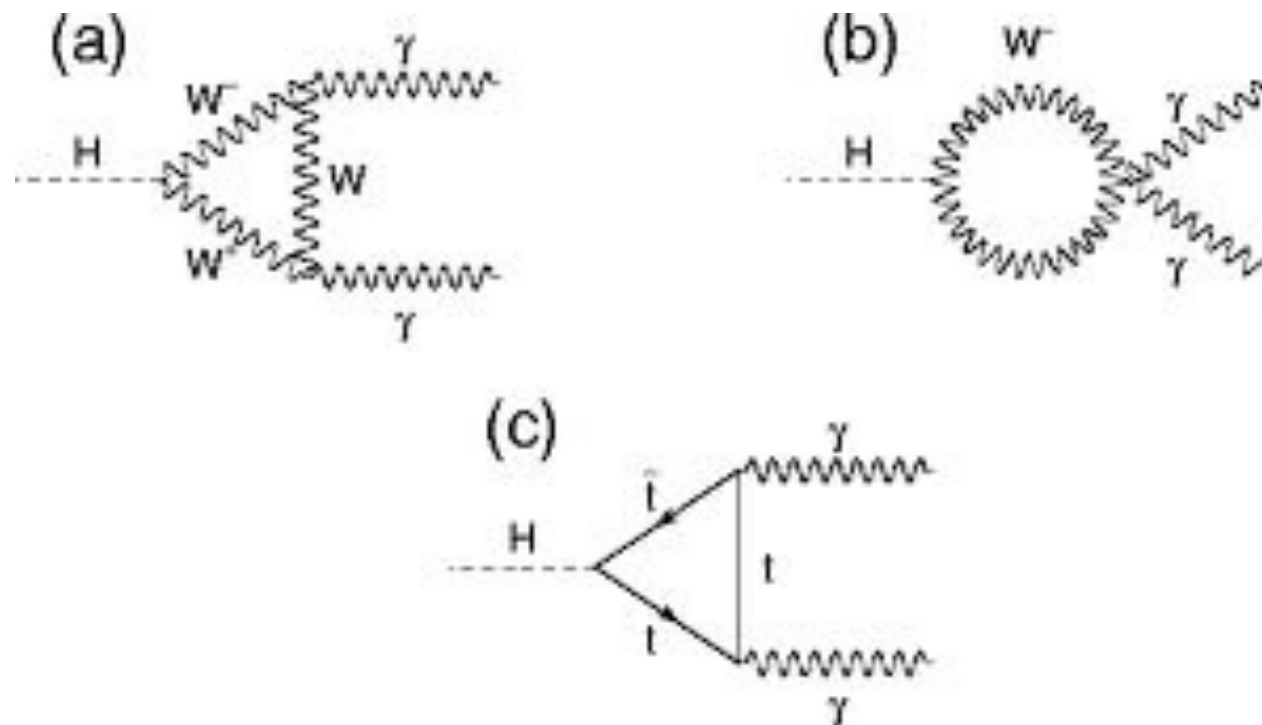
But, the ZZ signal rate tends to be enhanced as well - seems not very consistent with the current observation at the LHC





Enhanced $\gamma\gamma$ Decay Width ?

$$\sigma_{\gamma\gamma} = \sigma_{h_{\text{SM}}} \times \frac{\Gamma_{\gamma\gamma}}{\Gamma_{b\bar{b}} + \sum_{\text{SM}}^{i \neq b\bar{b}} \Gamma_i}$$



- ❏ Loop effect, mediated by charged particles
- ❏ Described by an effective theory

$$\mathcal{L}_{\text{eff}} = \frac{-\alpha_{\text{EM}} I}{2\pi} \frac{h_{\text{SM}}}{v_{\text{EW}}} F_{\mu\nu} F^{\mu\nu}$$



Effective Coupling

- Any particles coupled with the Higgs boson can get a mass from the Higgs VEV.
- I can be calculated through the photon self-energy corrections (e.g., M. Carena, et. al., 12')

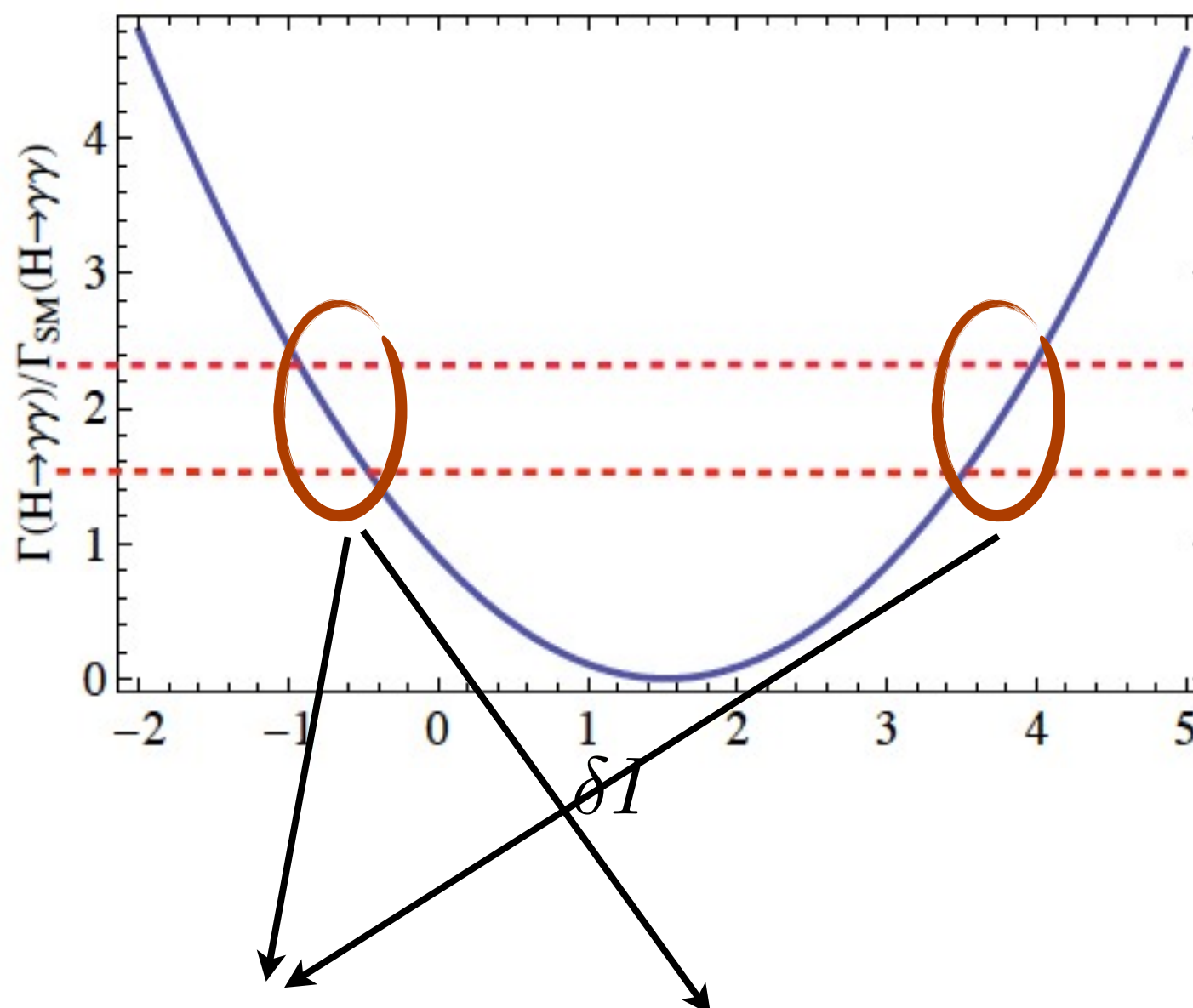
$$I = \sum_k \frac{b_k}{4} \frac{\partial}{\partial \log v_{EW}} \log (\det \mathcal{M}_k^2)$$

- SM - mainly controlled by the W and top-mediated loops: $I_W \sim -2.1$, $I_t \sim 0.5$
- Generalized to SUSY

$$I = \sum_k \frac{b_k}{4} \left[\cos \alpha \frac{\partial}{\partial v_u} \log (\det \mathcal{M}_k^2) - \sin \alpha \frac{\partial}{\partial v_d} \log (\det \mathcal{M}_i^2) \right]$$



Constructive Correction To Eff Coupling



There exist two solutions.

The more economical one is to get a negative δI , because $I_w + I_t$ is negative and such a NP correction is constructive.



Still Not Very Easy - Charged Fermion

- Yukawa-type mass only doesn't help much, because the sign of δI is fixed to be positive (similar to the contribution by top quark)

$$\delta I \propto \frac{Y_{f_i}}{M_{f_i}} = \frac{1}{v_{EW}}$$

- Extra mass sources are needed for the mediator. Consider a system, where the mediators also get mass from mixing

$$\mathcal{M}_f^\dagger \mathcal{M}_f = \begin{pmatrix} m_{11}^2 & m_{12}^2 \\ m_{12}^{*2} & m_{22}^2 \end{pmatrix}$$

$$\delta I \propto \frac{\alpha b_{1/2}}{16\pi (m_{11}^2 m_{22}^2 - |m_{12}^2|^2)} \left(m_{11}^2 \frac{\partial}{\partial v} m_{22}^2 + m_{22}^2 \frac{\partial}{\partial v} m_{11}^2 - \frac{\partial}{\partial v} |m_{12}^2|^2 \right)$$

- Economical choice: the off-diagonal ones contain Higgs VEV contributions while the diagonal ones are controlled by, e.g., vector-like or softly SUSY breaking corrections



Still Not Very Easy - Charged Fermion

- ❏ In supersymmetry, charginos provide such a possibility

$$M_{\chi^c}(h) = \begin{pmatrix} M_2 & gh_u \\ gh_d & \mu \end{pmatrix}$$

- ❏ But, the partial of the off-diagonal ones is controlled by the $SU(2)$ gauge coupling, which turns out to be not large enough.



Still Not Very Easy - Charged Scalar

- Single mass source can work. The simplest case is probably $Y_s H^\dagger H S^\dagger S$ with $Y_s < 0$, which gives negative correction.
- In SUSY, this type of interaction arises from the FF^* coupling of sfermions which is positive.
- Consider a system, where the mediators also get mass from mixing

$$\mathcal{M}_S^2 = \begin{pmatrix} \tilde{m}_L(v)^2 & \frac{1}{\sqrt{2}}vX_S \\ \frac{1}{\sqrt{2}}vX_S & \tilde{m}_R(v)^2 \end{pmatrix}$$

$$\frac{\partial \log(\det \mathcal{M}_S^2)}{\partial v} \simeq v \frac{(m_{L0}^2 + \frac{1}{2}c_L v^2)c_R + (m_{R0}^2 + \frac{1}{2}c_R v^2)c_L - X_S^2}{m_{S_1}^2 m_{S_2}^2}.$$

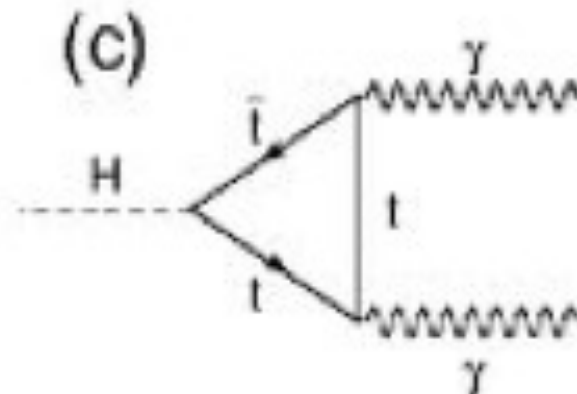
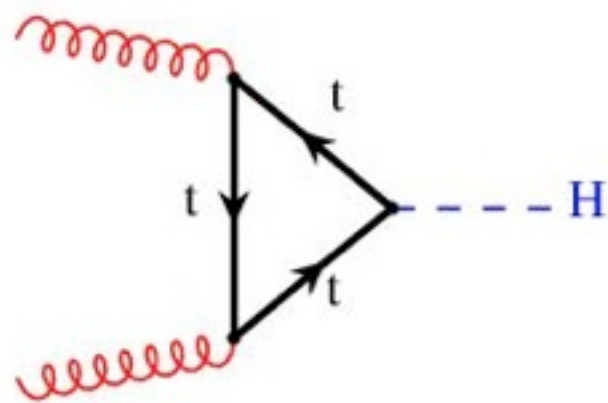
$$\tilde{m}_L^2 = \tilde{m}_{L0}^2 + \frac{1}{2}c_L v^2, \quad \tilde{m}_R^2 = \tilde{m}_{R0}^2 + \frac{1}{2}c_R v^2,$$

- Economical choice for SUSY: relatively large A -term + relatively light scalar



Still Not Very Easy - Charged Scalar

- ☒ In supersymmetry, stop quarks indeed can enhance the di-photon width in such a way, but not very helpful for the signal rate enhancement



- ☒ But, stau leptons do work, although marginal [Carena, et. al., 2011]



Why Extensions Of The MSSM ?

- What we learn from the SM: the Higgs mass is mainly controlled by the quartic coupling in the Higgs potential
- In the MSSM,
 - quartic terms arise from the SM D-terms

$$m_h^2 = m_Z^2 \cos^2 2\beta + \text{loop} \quad \text{loop} \propto \log \left(\frac{M_{\text{SUSY}}}{M_{\text{top}}} \right)$$

- $m_h = 125 \text{ GeV}$ needs significant loop-level corrections, requiring

$$\Lambda_{\text{SUSY}} \gg m_{\text{top}}$$

- To lift the Higgs mass at tree-level, new quartic terms are needed



Extra U(1) Gauge Symmetry

- ❏ One possibility is from new non-decoupling D-term
- ❏ We will focus on gauged Peccei-Quinn symmetry:
 - ❏ Higgs fields are charged by definition. Bare μ term is forbidden
 - ❏ Provides a solution to the μ problem in the MSSM

$$W \sim \lambda S H_u H_d, \text{ with } \mu_{\text{eff}} = \lambda \langle S \rangle$$

- ❏ It is anomalous. Charged anomaly spectators are required
- ❏ Good! If some charged exotics happen to be light, they may help implement the mechanisms of the di-photon decay enhancement



Effective Theory

- U(1)_{PQ} breaking can be quite involved. We focus on a simplified scenario.
 - PQ symmetry breaking scale $f_{PQ} > M_{Z'} > \text{EW scale}$
 - Integrate out the radial modes. Keep saxion only

SSB by \mathbf{S}_i , then we have

$$\mathbf{S}_i = f_i e^{q_i \mathbf{A}} / f_{PQ} \quad f_{PQ} = \sqrt{\sum_i q_i^2 f_i^2}$$

$$\mathbf{A} = A + \sqrt{2}\theta\tilde{a} + \theta^2 F_A, \quad A = \frac{1}{\sqrt{2}}(s + ia)$$

$$\mathbf{W}_H = \lambda \mathbf{S} \mathbf{H}_u \mathbf{H}_d = \lambda f_S e^{q_S \mathbf{A}} / f_{PQ} \mathbf{H}_u \mathbf{H}_d,$$

$$\mathbf{K} = \sum_i f_i^2 \exp\left(\frac{q_i(\mathbf{A} + \mathbf{A}^\dagger)}{f_{PQ}} + 2g_{PQ}q_i \mathbf{V}_{PQ}\right) + \sum_a \mathbf{H}_a^\dagger \exp(2g_{PQ}q_a \mathbf{V}_{PQ} + \mathbf{U}_{SM}) \mathbf{H}_a,$$



Effective Theory

Further integrating out the saxion in the PQ sector, we have

$$\begin{aligned}
 V_{WZ} = & (|\mu_{\text{eff}}|^2 + m_{H_u}^2)|H_u|^2 + (|\mu_{\text{eff}}|^2 + m_{H_d}^2)|H_d|^2 \\
 & - 2B_\mu \text{Re}(H_u H_d) + \frac{1}{8}(g_2^2 + g_Y^2)(|H_u|^2 - |H_d|^2)^2 \\
 & - g_{PQ} q_{H_u} \langle D_{PQ} \rangle (|H_u|^2 + |H_d|^2) \\
 & + a_1 |H_u H_d|^2 + a_2 (|H_u|^2 + |H_d|^2)^2 \\
 & + a_3 \text{Re}(H_u H_d) (|H_u|^2 + |H_d|^2) . \tag{4}
 \end{aligned}$$

$$a_1 = \mathcal{O}(\lambda^2) , \quad a_2 = \frac{1}{2} g_{PQ}^2 q_{H_u}^2 \delta^2 + \mathcal{O}(\lambda^2) ,$$

$$a_3 = \frac{-4A_\lambda \lambda g_{PQ}^2 q_{H_u}^2 f_S}{m_s^2} + \mathcal{O}(\lambda^3) .$$

$$m_s^2 = 2g_{PQ}^2 f_{PQ}^2 (1 + \delta^2) \text{ with } \delta^2 = \frac{\sum_i m_{S_i}^2 q_i^2 f_i^2}{g_{PQ}^2 f_{PQ}^4}$$

Non-trivial corrections to the MSSM potential due to the PQ D-term require sizable softly SUSY breaking effects in the PQ sector!



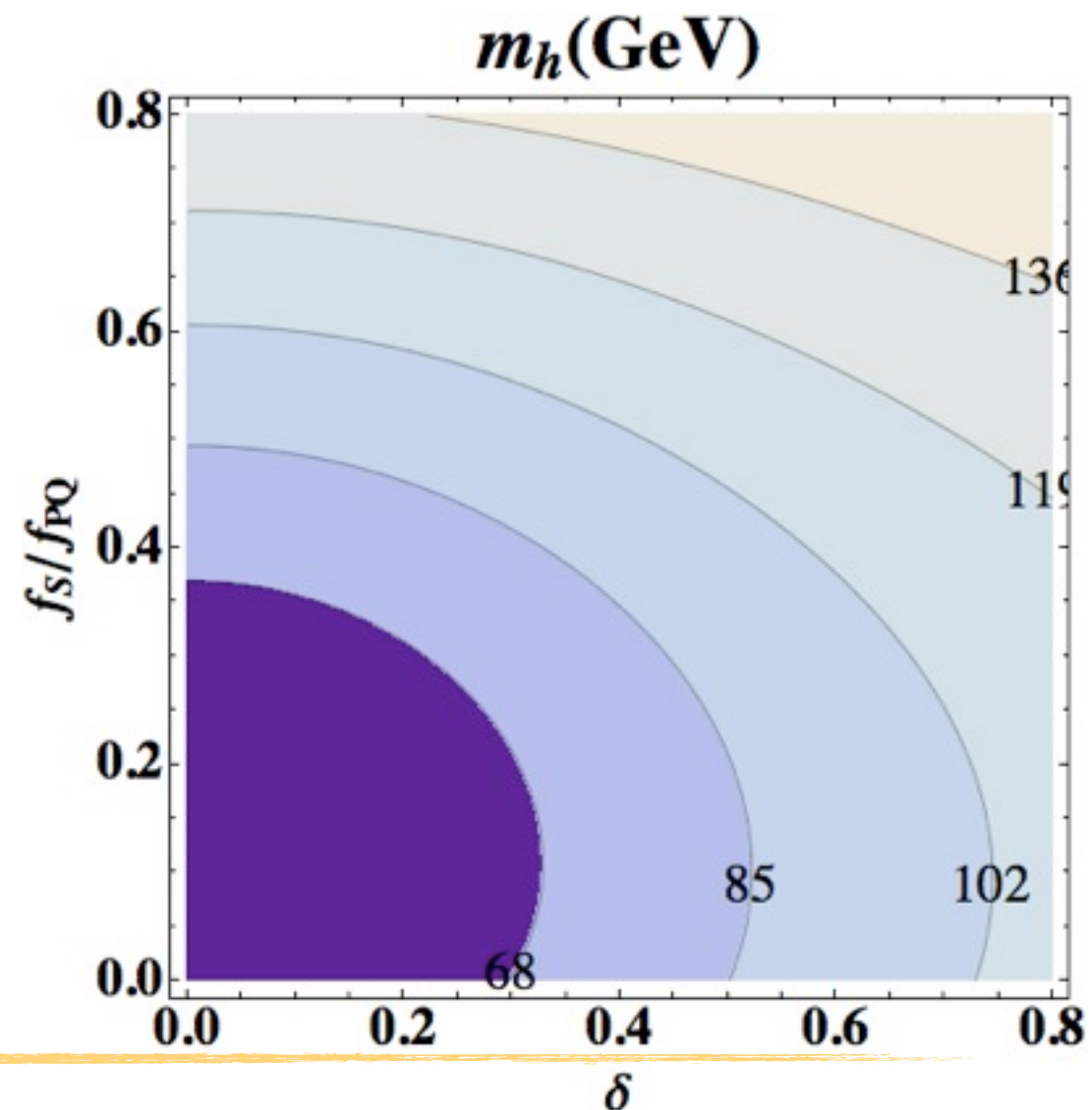
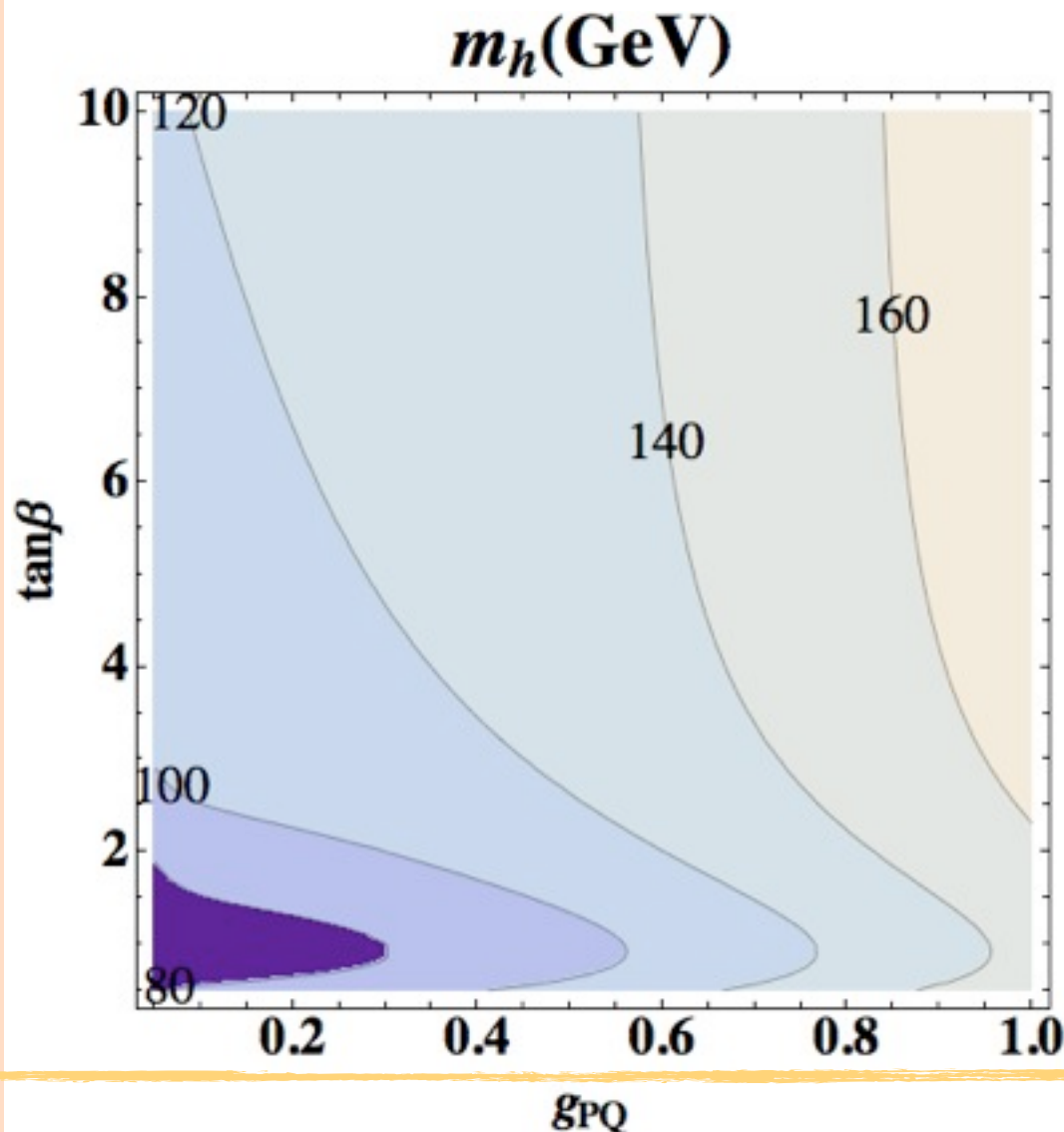
Gauged Peccei-Quinn Symmetry

$$(M_h^2)_{\text{tree}} \approx m_Z^2 \cos^2 2\beta + \left(\frac{a_1}{2} \sin^2 2\beta + 2a_2 + a_3 \sin 2\beta \right) v_{\text{EW}}^2$$

Six free parameters at tree level :

$$g_{\text{PQ}}, f_{\text{PQ}}, \tan \beta, \frac{f_S}{f_{\text{PQ}}}, \lambda, \frac{A_\lambda}{f_{\text{PQ}}}$$

↓
New contribution





One Anomaly-free U(1)PQ Model

Vector-like mass terms for non-colored exotic charged fermions

$$\begin{aligned}
 \mathbf{W} = & \mathbf{W}_H + \beta^{pq} \mathbf{S} \mathbf{D}_p \mathbf{D}_q^c \\
 & + \delta_N \mathbf{S}_1 \mathbf{N} \mathbf{N}^c + \delta_X \mathbf{S}_1 \mathbf{X} \mathbf{X}^c \\
 & + \gamma^p (\mathbf{H}_u \mathbf{D}_p \mathbf{X}^c + \mathbf{H}_d \mathbf{D}_p \mathbf{N}^c) \\
 & + \gamma_c^q (\mathbf{H}_d \mathbf{D}_q^c \mathbf{X} + \mathbf{H}_u \mathbf{D}_q^c \mathbf{N}) \\
 & + \mathbf{W}_Y (\mathbf{H}_u \leftrightarrow \mathbf{D}_k, \mathbf{H}_d \leftrightarrow \mathbf{D}_k^c) \\
 & + \mathbf{W}_{LQ} + \mathbf{W}_S
 \end{aligned}$$

Indeed, there exist color-neutral superfields which can couple to Higgs fields

Particles	Gauge charges	Particles	Gauge charges
\mathbf{L}_i	(1; 2; -1/2; 1/2)	\mathbf{Q}_i	(3; 2; 1/6; 1/2)
$\bar{\mathbf{N}}_i$	(1; 1; 0; 1/2)	$\bar{\mathbf{u}}_i$	(3; 1; -2/3; 1/2)
$\bar{\mathbf{e}}_i$	(1; 1; 1; 1/2)	\mathbf{d}_i	(3; 1; 1/3; 1/2)
\mathbf{H}_d	(1; 2; -1/2; -1)	\mathbf{H}_u	(1; 2; 1/2; -1)
\mathbf{T}_1	(3; 1; 2/3; -1)	\mathbf{T}_1^c	(3; 1; -2/3; -1)
\mathbf{T}_2	(3; 1; 2/3; -1)	\mathbf{T}_2^c	(3; 1; -2/3; -1)
\mathbf{T}_3	(3; 1; -1/3; -1)	\mathbf{T}_3^c	(3; 1; 1/3; -1)
\mathbf{D}_1	(1; 2; 1/2; -1)	\mathbf{D}_1^c	(1; 2; -1/2; -1)
\mathbf{D}_2	(1; 2; 1/2; -1)	\mathbf{D}_2^c	(1; 2; -1/2; -1)
\mathbf{X}	(1; 1; 1; 2)	\mathbf{X}^c	(1; 1; -1; 2)
\mathbf{N}	(1; 1; 0; 2)	\mathbf{N}^c	(1; 1; 0; 2)
\mathbf{S}	(1; 1; 0; 2)	\mathbf{S}^c	(1; 1; 0; -2)
\mathbf{S}_1	(1; 1; 0; -4)	\mathbf{S}_1^c	(1; 1; 0; 4)
\mathbf{S}_2	(1; 1; 0; -2)		

$$\mathbf{W}_{LQ} = \alpha^{pq} \mathbf{S} \mathbf{T}_p \mathbf{T}_q^c + \alpha^{33} \mathbf{S} \mathbf{T}_3 \mathbf{T}_3^c + \kappa^{rs} \mathbf{L}_r \mathbf{Q}_s \mathbf{T}_3 + \kappa^{rsp} \bar{\mathbf{N}}_r \bar{\mathbf{u}}_s \mathbf{T}_p$$



One Benchmark: Light Charged Exotic Fermion

g_{PQ}	f_{PQ} (GeV)	f_S/f_{PQ}	A_λ/f_{PQ}	λ
0.6	2500	0.4	0.1	0.3
$\tan \beta$	δ	A_γ (GeV)	A_{γ_c} (GeV)	γ, γ_c
1.3	0.6	0	0	1.6
m_D (GeV)	m_X (GeV)	$m_{\tilde{D}, \tilde{X}, \tilde{N}}^2$ (GeV ²)	$A_{\tilde{t}}$ (GeV)	$m_{\tilde{Q}_3, \tilde{t}}^2$ (GeV ²)
440	330	1000 ²	1200	500 ²
a_1	a_2	a_3	B_μ (GeV ²)	μ_{eff} (GeV)
0.06	0.9	-0.02	7.5×10^4	300
m_h (GeV)	$m_{\psi_{1^c}}$ (GeV)	$m_{\psi_{1^0}}$ (GeV)	$m_{\phi_{1^c}}$ (GeV)	$m_{\phi_{1^0}}$ (GeV)
125	105	105	943	943
$R(h \rightarrow \gamma\gamma)$	ΔS	ΔT		
1.8	0.11	0.10		

Enhanced di-photon signal rate!

Relatively small DeltaT, because we work in the region with approximate custodial symmetry

Unlike charginos in the MSSM, the coupling between h and the charged light exotic fermion is controlled by Yukawa couplings



One Benchmark: Light Charged Exotic Scalar

g_{PQ}	f_{PQ} (GeV)	f_S/f_{PQ}	A_λ/f_{PQ}	λ
0.6	2500	0.4	0.1	0.3
$\tan \beta$	δ	A_γ (GeV)	A_{γ_c} (GeV)	γ, γ_c
6	0.6	1440	1000	0.5
m_D (GeV)	m_X (GeV)	$m_{\tilde{D}, \tilde{X}, N}^2$ (GeV ²)	$A_{\tilde{t}}$ (GeV)	$m_{\tilde{Q}_3}^2, m_{\tilde{t}}^2$ (GeV ²)
500	350	100^2	1200	500^2
a_1	a_2	a_3	B_μ (GeV ²)	μ_{eff} (GeV)
0.06	0.07	-0.02	7.5×10^4	300
m_h (GeV)	$m_{\psi 1^c}$ (GeV)	$m_{\psi 1^0}$ (GeV)	$m_{\phi 1^c}$ (GeV)	$m_{\phi 1^0}$ (GeV)
125	325	325	104	233
$R(h \rightarrow \gamma\gamma)$	ΔS	ΔT		
1.7	0.03	0.08		

Enhanced di-photon signal rate!

Leading to a large coupling between h and the charged light exotic Scalar! Recall - it is controlled by

$$H_u D_p X^c$$



Conclusions

- ❧ The two benchmarks are presented in a specific model, but they represent large classes of models in which an enhancement of the $h \rightarrow \gamma\gamma$ signal rate does not lead to a violation of the EWPT constraints.
- ❧ Benchmark I: light charged fermion + large coupling with Higgs field + approximate custodial symmetry
- ❧ Benchmark II: light charged scalar + relatively large A -parameter
- ❧ If only the $h \rightarrow \gamma\gamma$ signal rate is enhanced while the other ones are not modified, it is very likely that the new exotics carry the EW charges only
- ❧ The collider signals of their directly search are similar to the EW-ino and the slepton ones



Thank you!